



Physics and Political Update

NOvA Collaboration Meeting

Fermilab

2 October 2004

Gary Feldman



Topics

- **Presentation to the PAC Aspen Meeting**
- **PAC Response**
- **Visit to DoE**
- **Outlook**



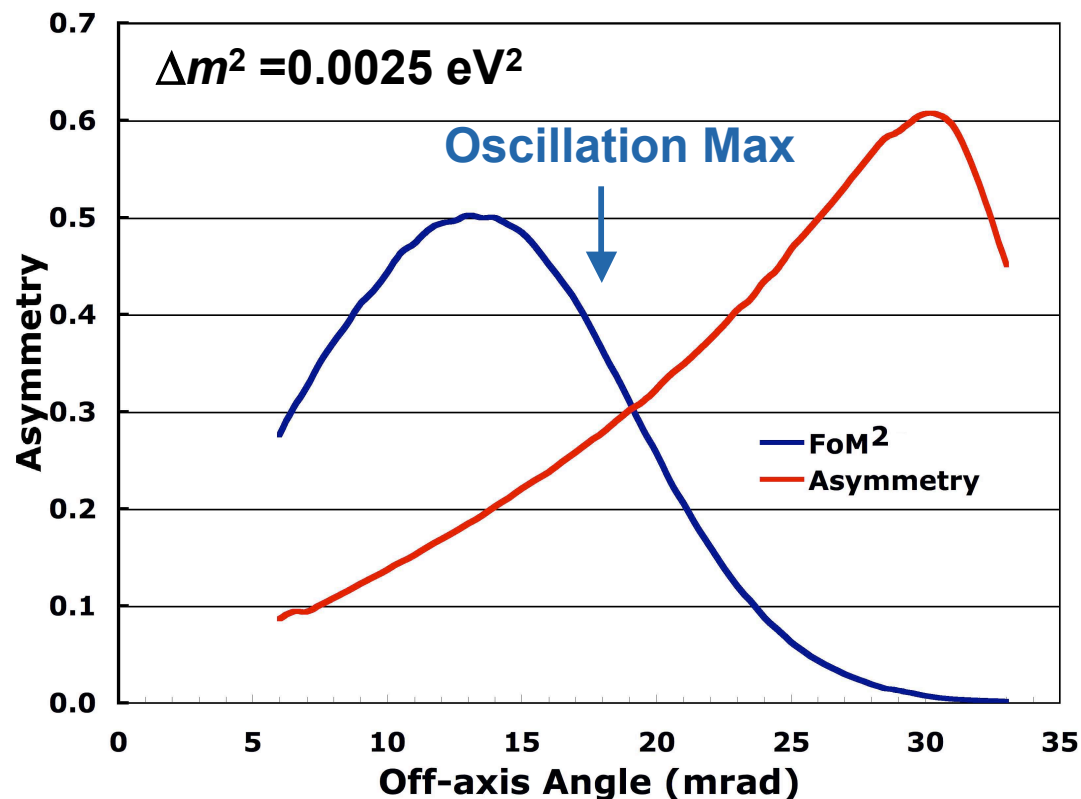
PAC Concerns from the April Meeting

- The PAC seemed to be concerned that
 - NOvA is just a duplication of T2K, and if T2K runs earlier, NOvA could be irrelevant, and that
 - NOvA is being optimized for the near term and will not be the right detector for later use.
- We submitted Appendix D to the June meeting
 - to reorient NOvA toward its unique role in resolving the mass hierarchy,
 - to show that there is a progression of steps that allows the resolution of the mass hierarchy for all values of the CP phase δ and an order of magnitude range of $\sin^2(2\theta_{13})$, and
 - to show that NOvA is optimized for all stages of this progression, even with reasonable uncertainty on the value of Δm^2 .



The Optimization Problem

FoM² and Asymmetry vs. Angle



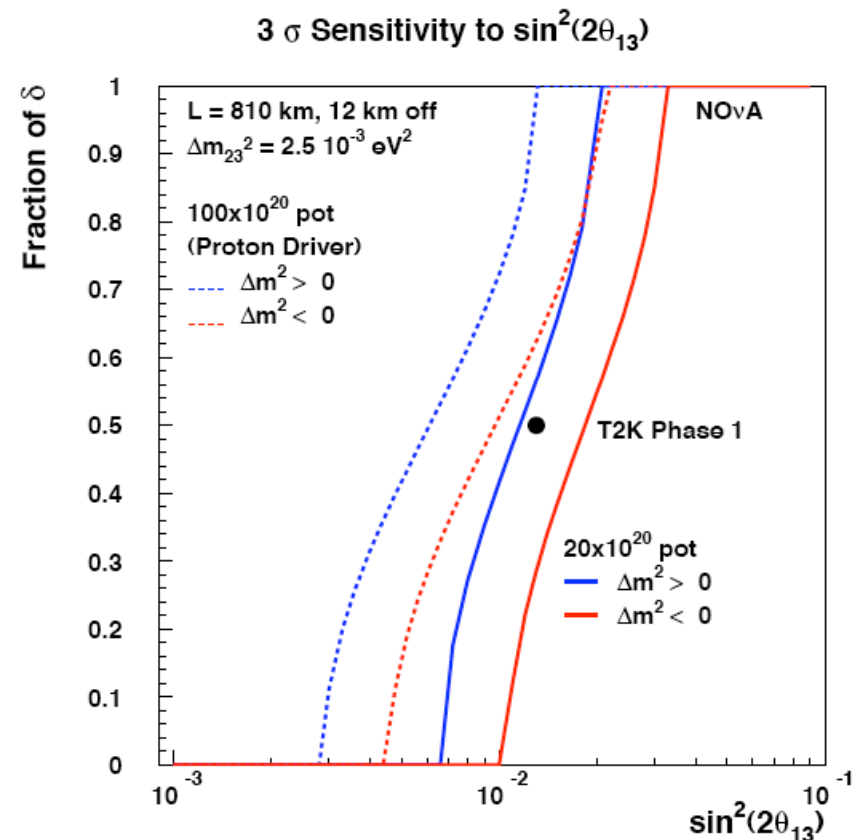
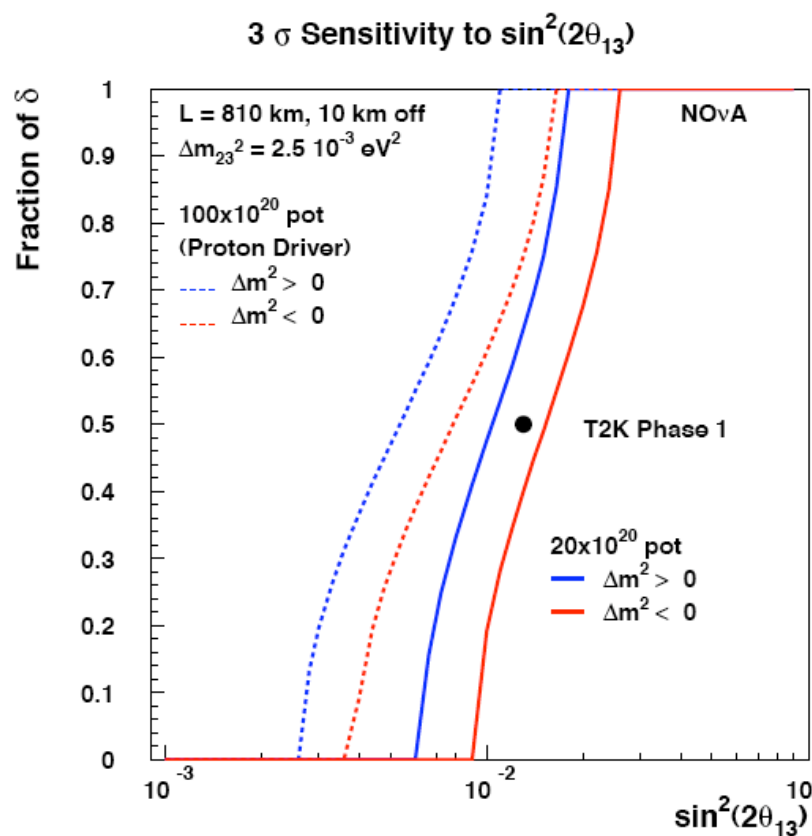
$$\text{FoM} = \frac{\text{signal}}{\sqrt{\text{background}}}$$

Asymmetry =

$$\left(\frac{\sigma_{\nu} - \sigma_{\bar{\nu}}}{\sigma_{\nu} + \sigma_{\bar{\nu}}} \right)$$

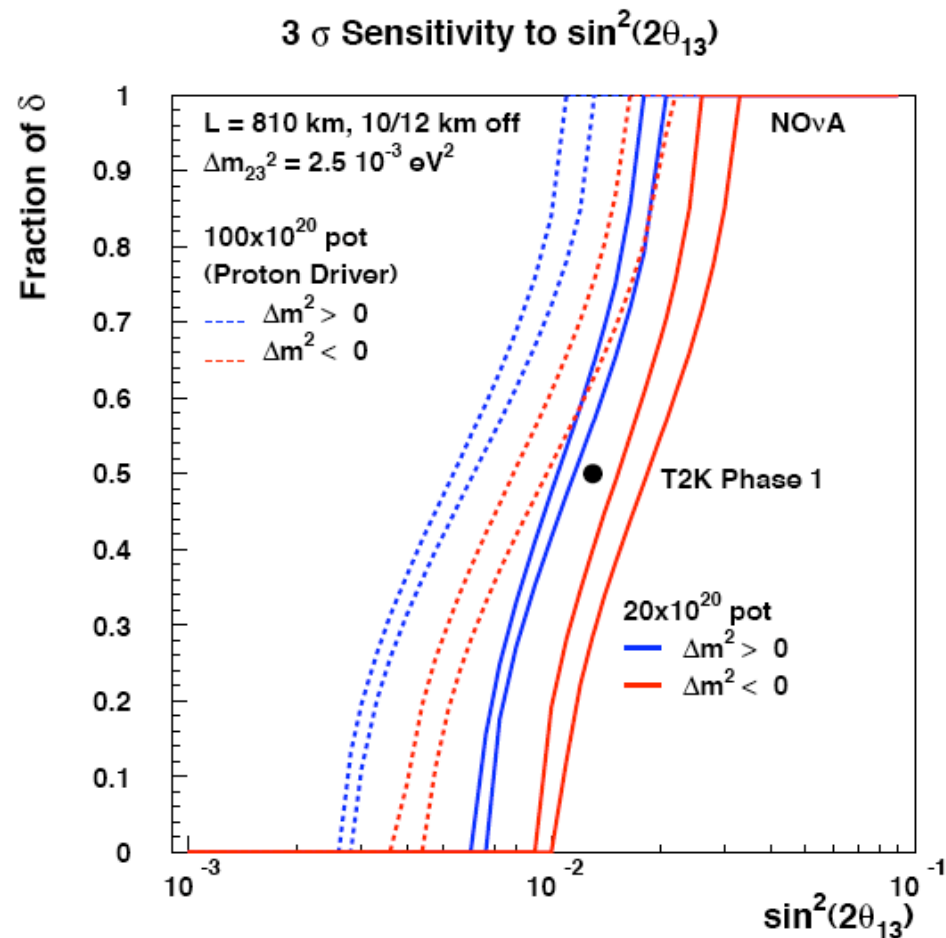


3 σ Discovery Potential for $\nu_\mu \rightarrow \nu_e$



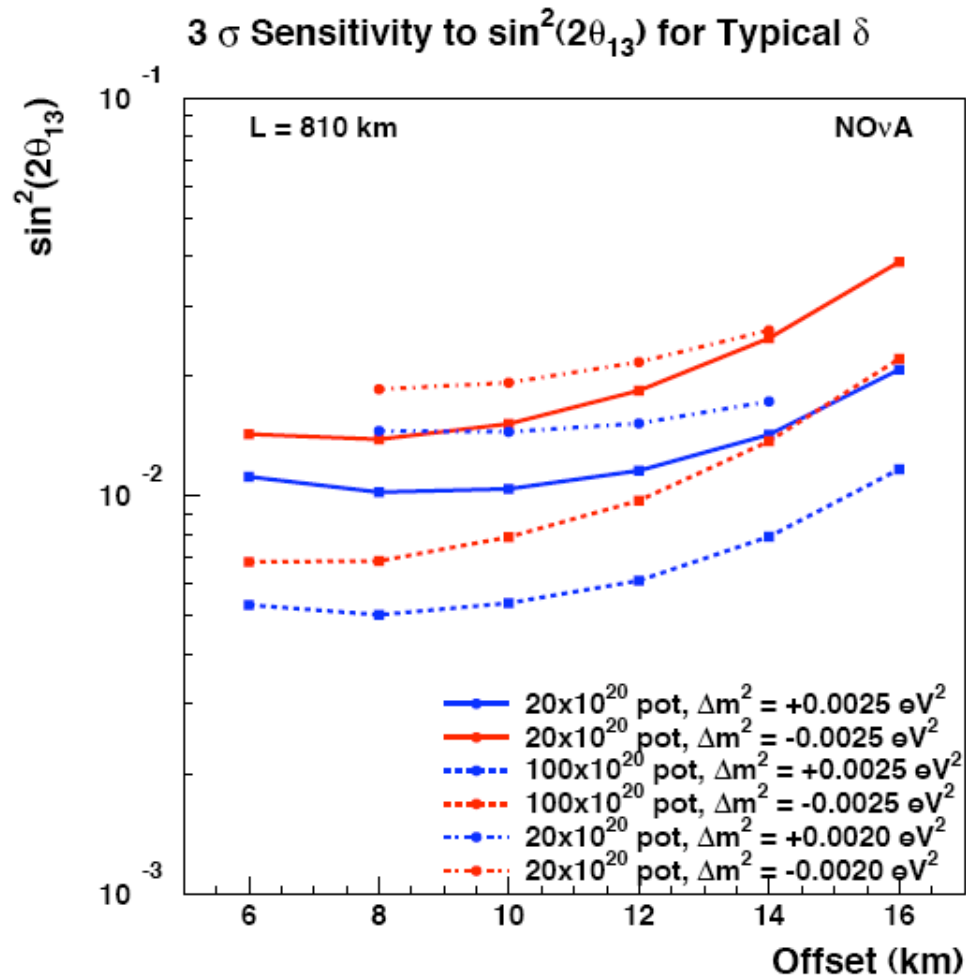


Comparison of 10 and 12 km





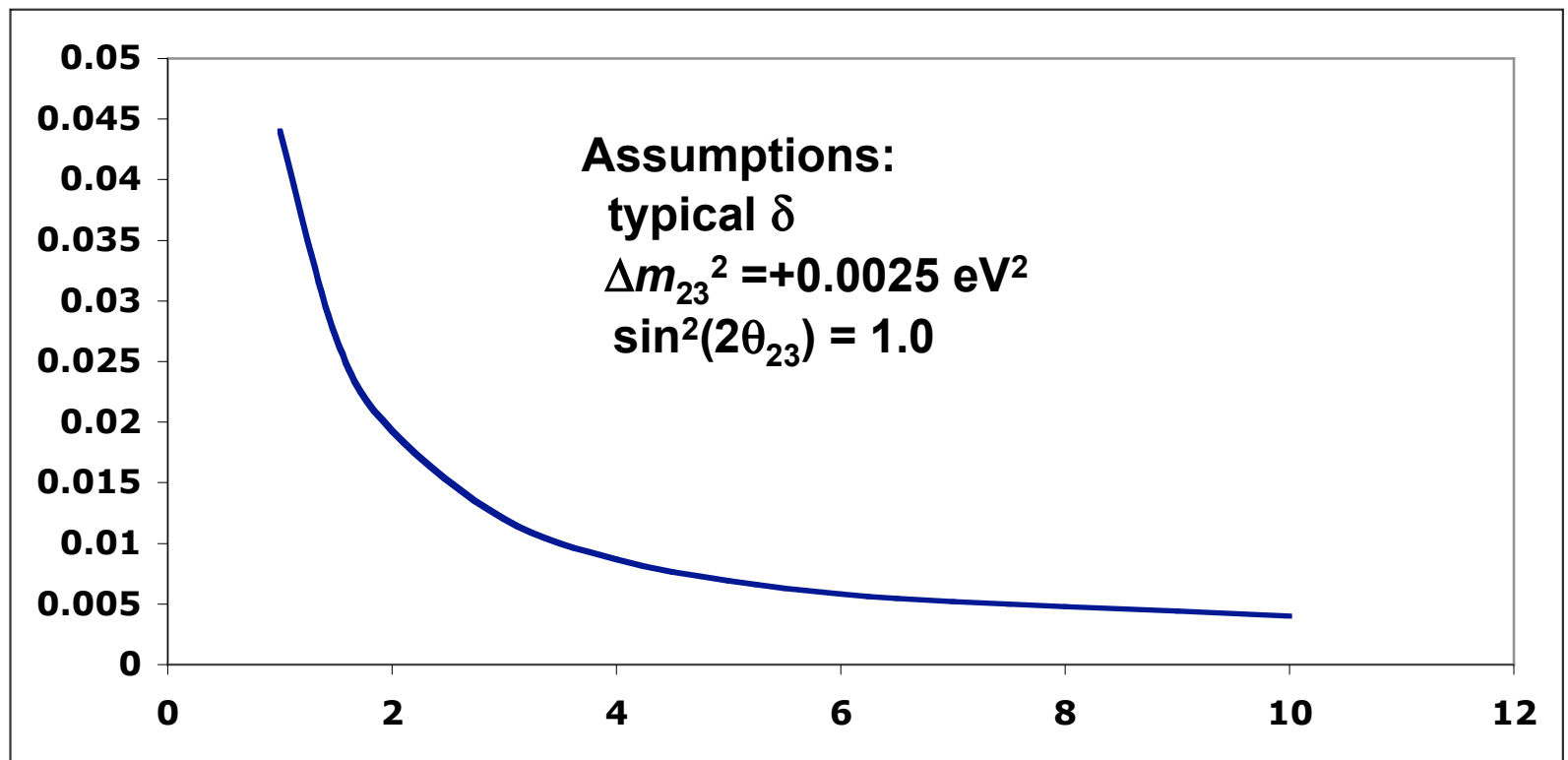
3 σ Discovery Potential for $\nu_\mu \rightarrow \nu_e$ vs. Off-Axis Distance



Note: There is a loss of sensitivity for $\Delta m^2 = 0.002$ eV², but not a loss of range, since the CHOOZ limit is correspondingly weaker there.



95% CL on $\sin^2(2\theta_{13})$ vs. years from start of installation



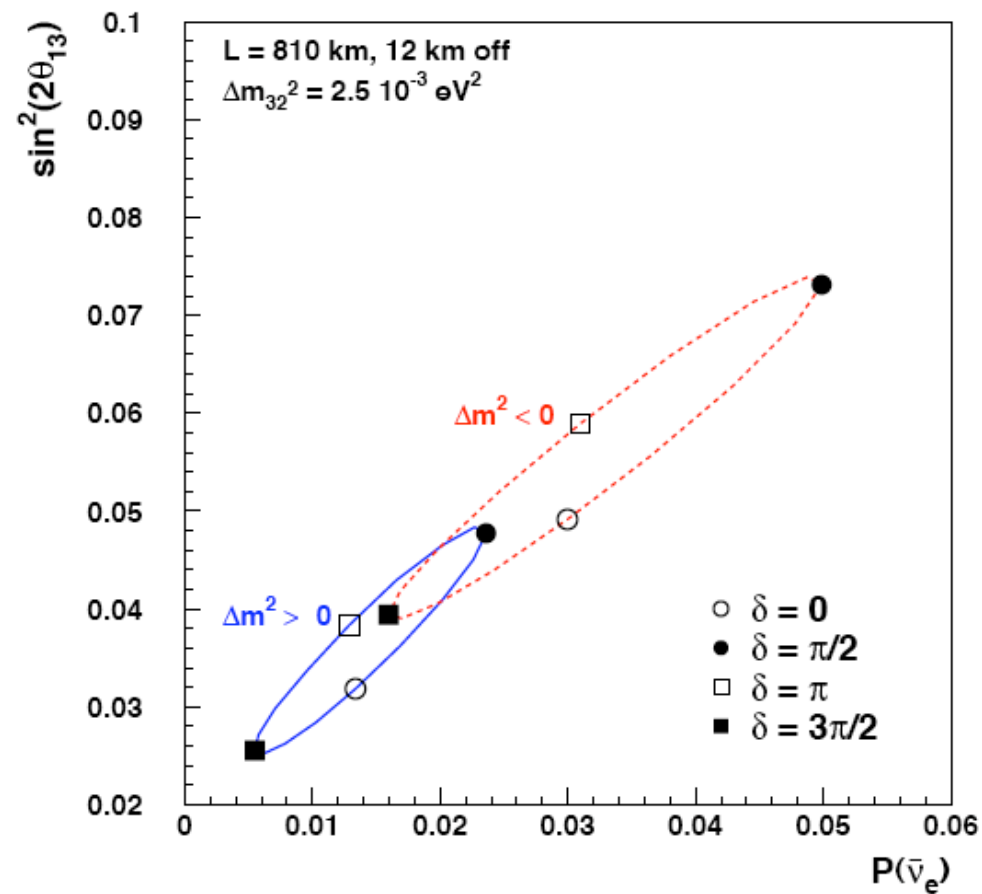
Years from start of installation



Reminder of the Problem

Part 1

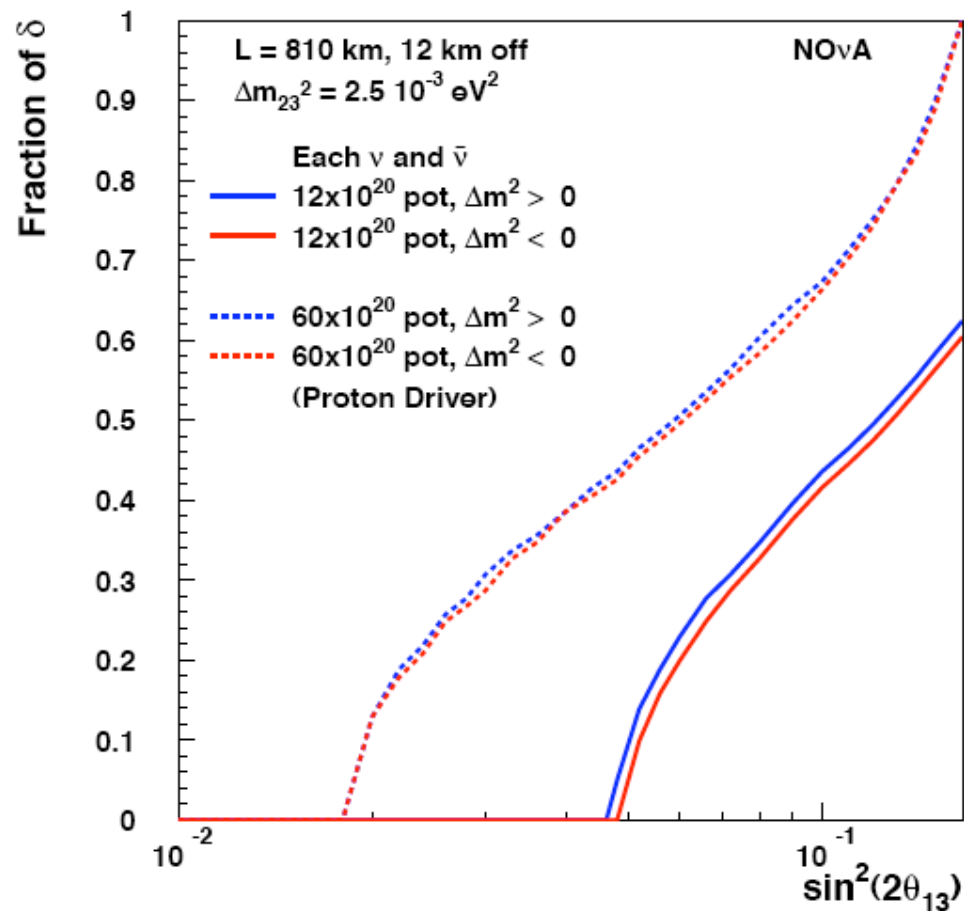
$\sin^2(2\theta_{13})$ vs. $P(\bar{\nu}_e)$ for $P(\nu_e) = 0.02$





95% CL Resolution of the Mass Hierarchy

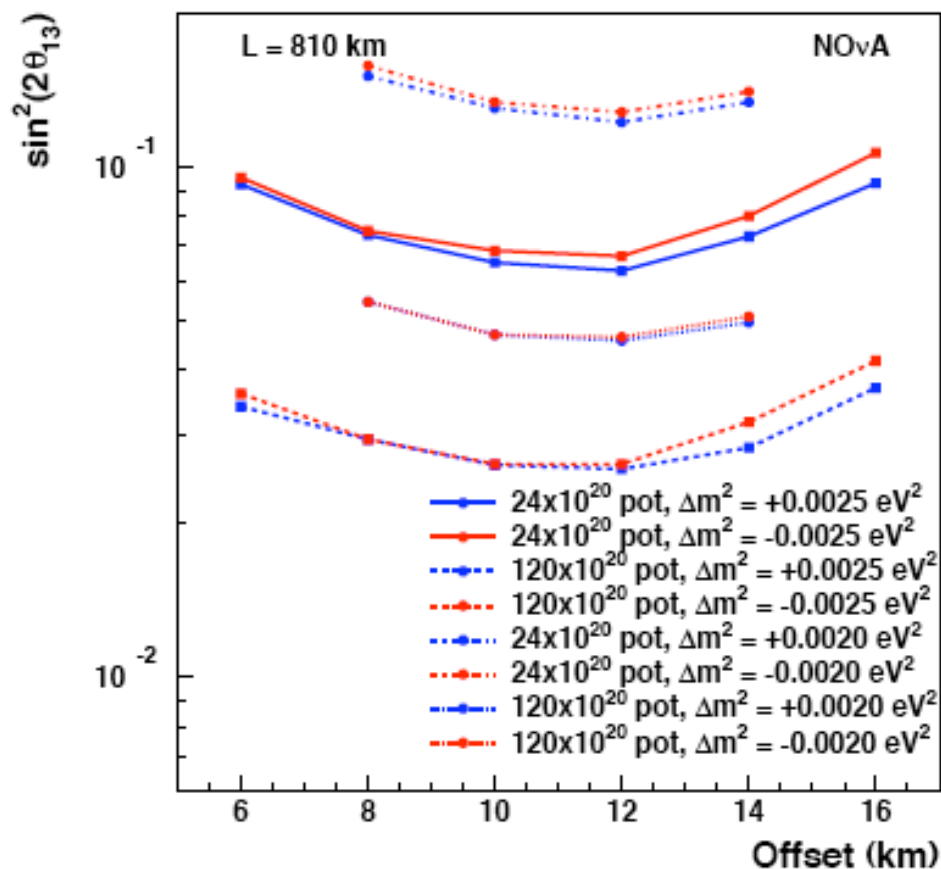
2 σ Resolution of the Mass Hierarchy





Mass Hierarchy Resolution vs. Off-Axis Distance

2σ Mass Hierarchy Resolution for 1st Quartile δ

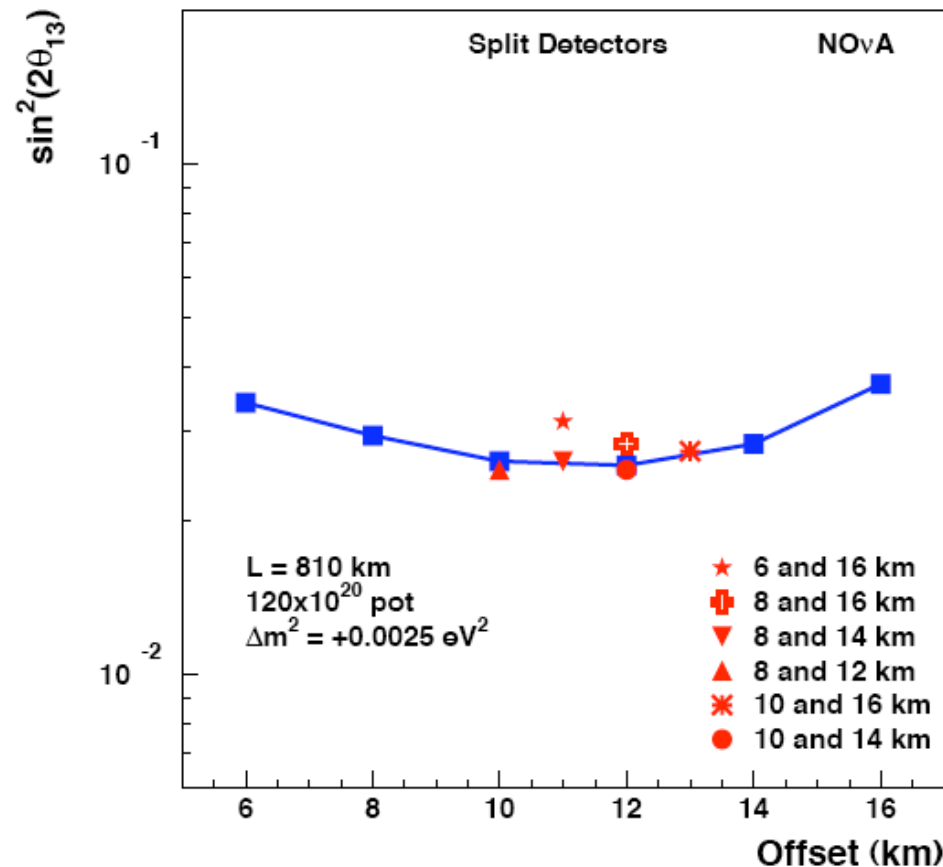


**12 km off-axis
is best for both
 $\Delta m^2 = 0.0025$ and
 $\Delta m^2 = 0.0020$ eV²**



PAC Question: Are two (1/2) Detectors Better than One?

2 σ Mass Hierarchy Resolution for 1st Quartile δ

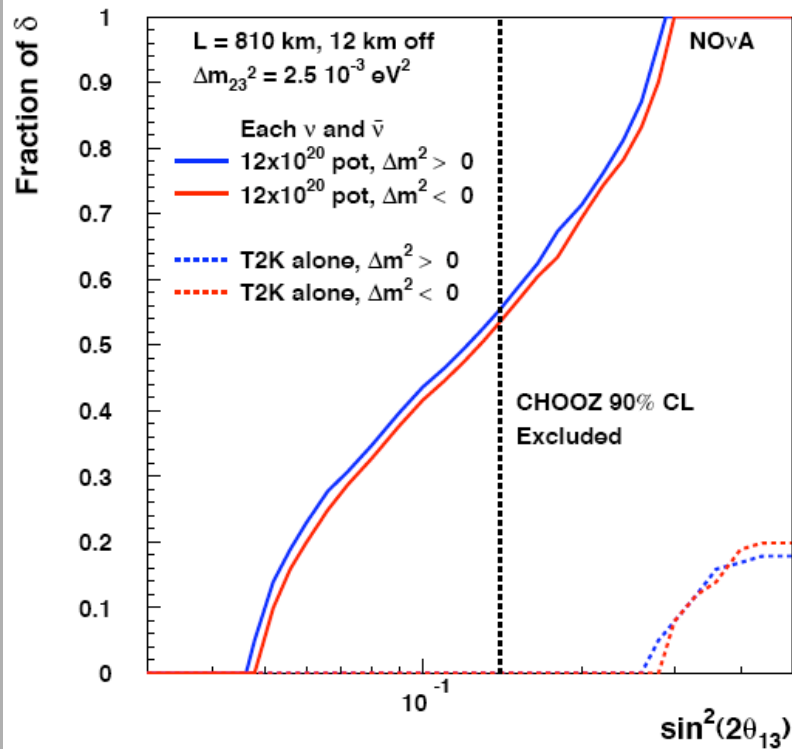


Answer: Yes, but not by enough to overcome the fiducial and infrastructure costs.



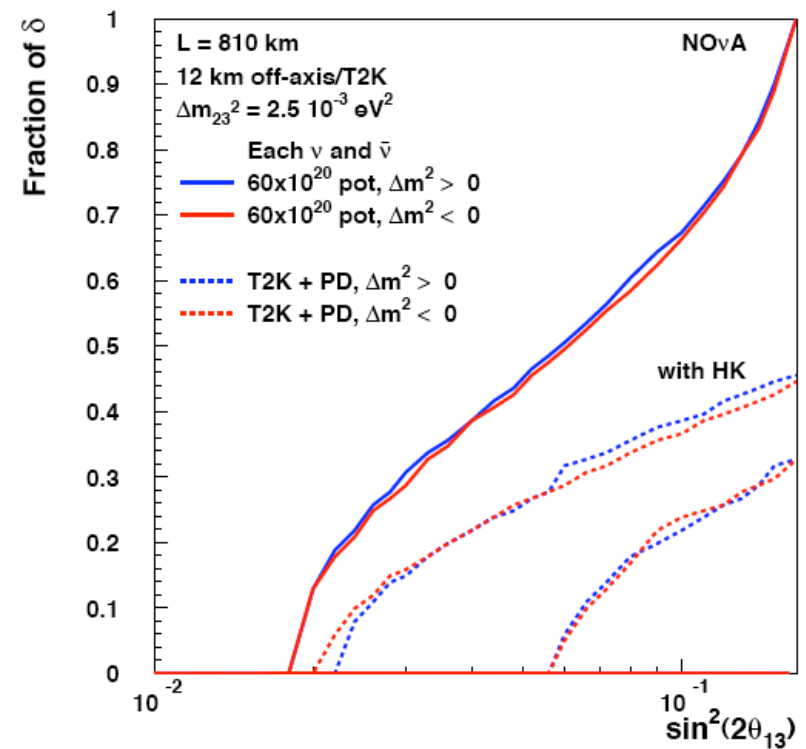
NOvA Alone vs. T2K Alone

2 σ Resolution of the Mass Hierarchy



Note change of horizontal scale

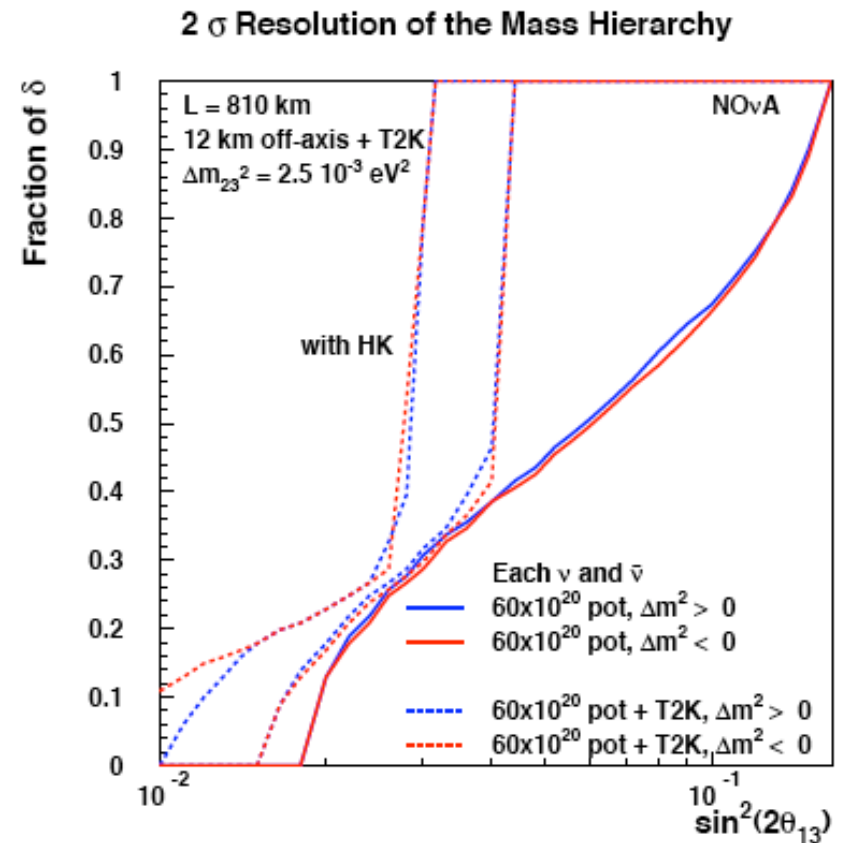
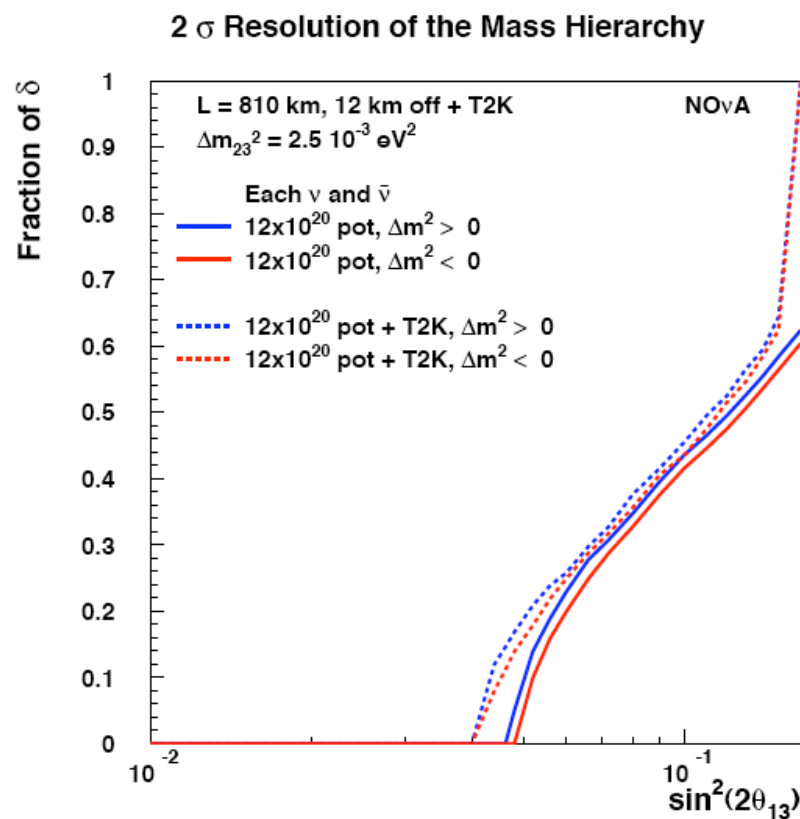
2 σ Resolution of the Mass Hierarchy



Proton Drivers



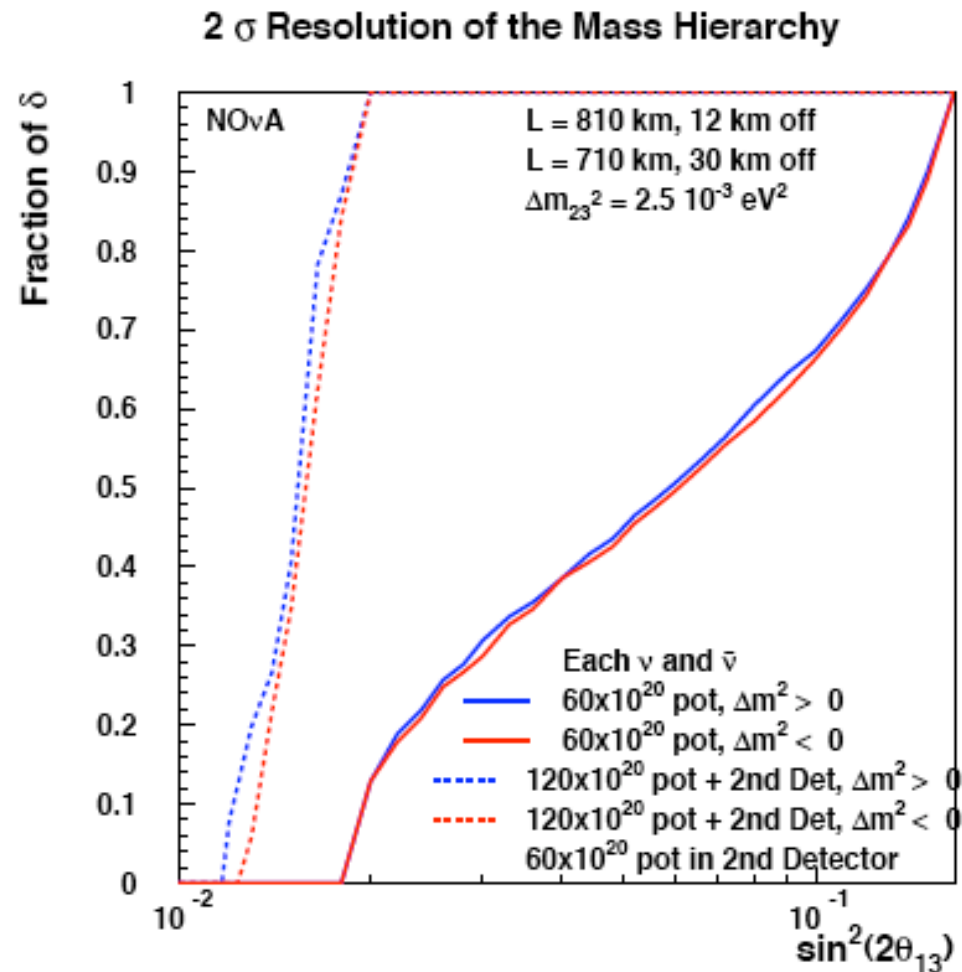
Combination with T2K



Proton Drivers



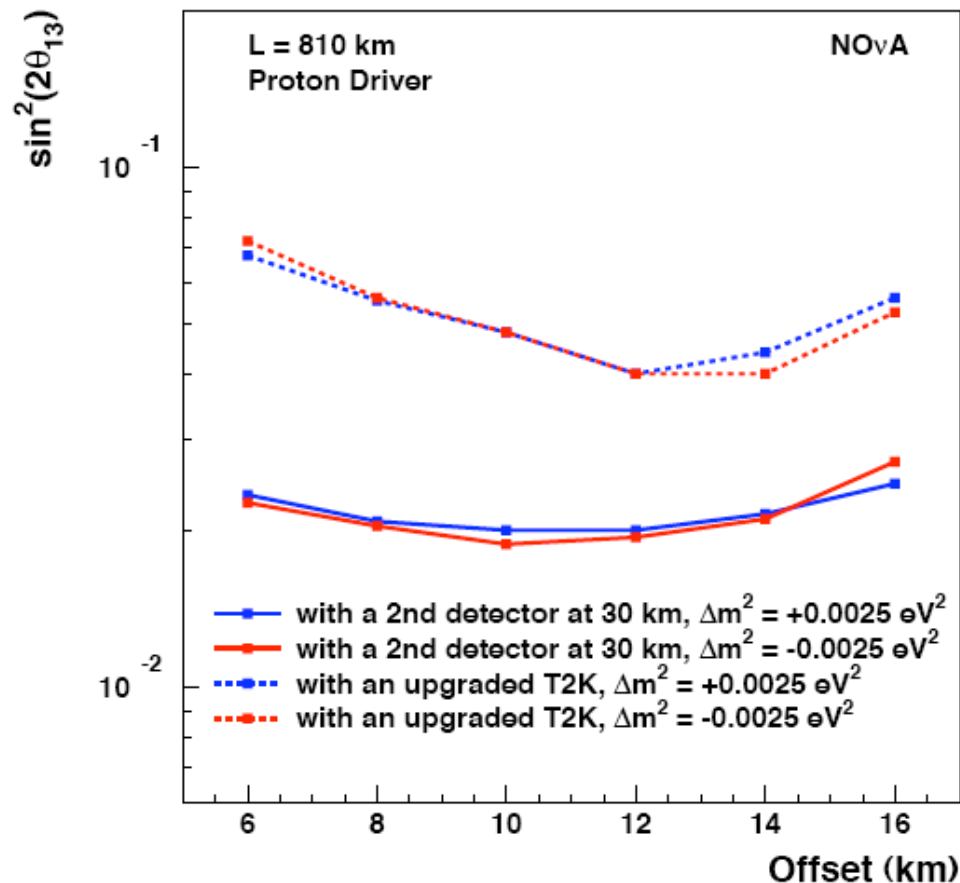
Combination with a 2nd OA Detector at the 2nd Maximum





Mass Hierarchy Resolution vs. Off-Axis Distance

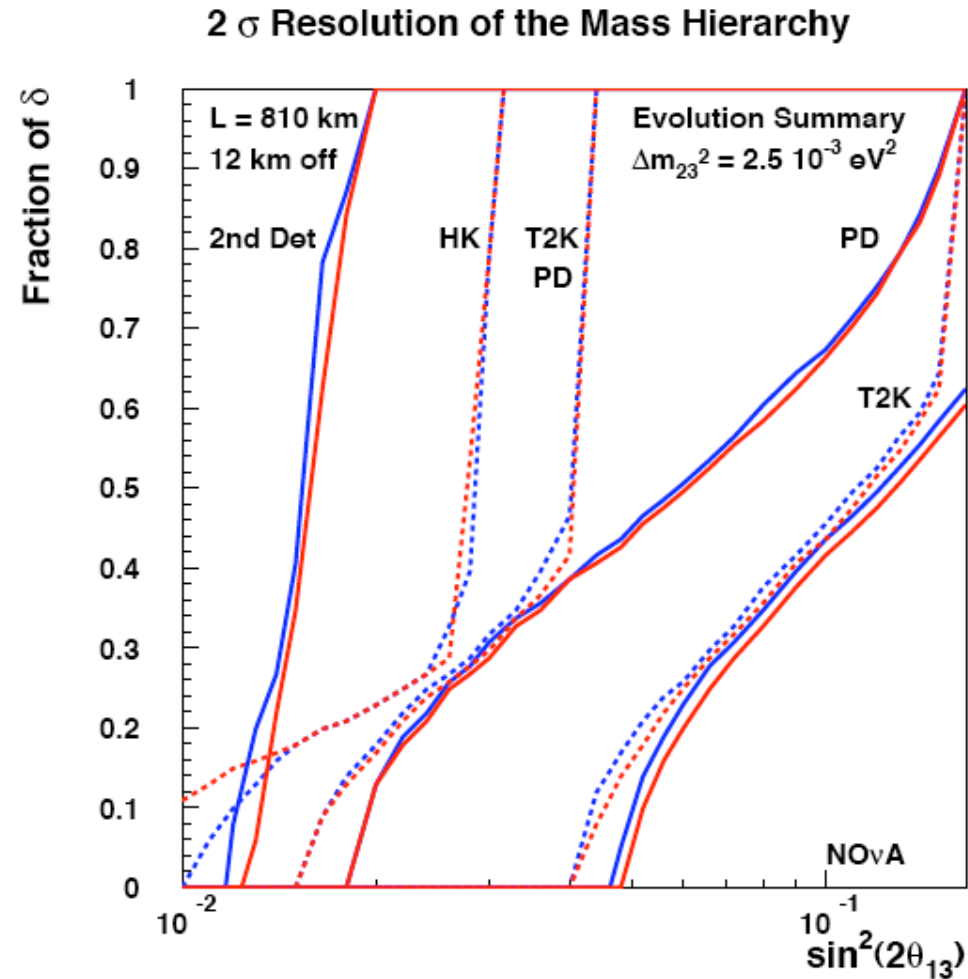
2σ Mass Hierarchy Resolution for all δ



**Again, 12 km
provides a good
optimization.**



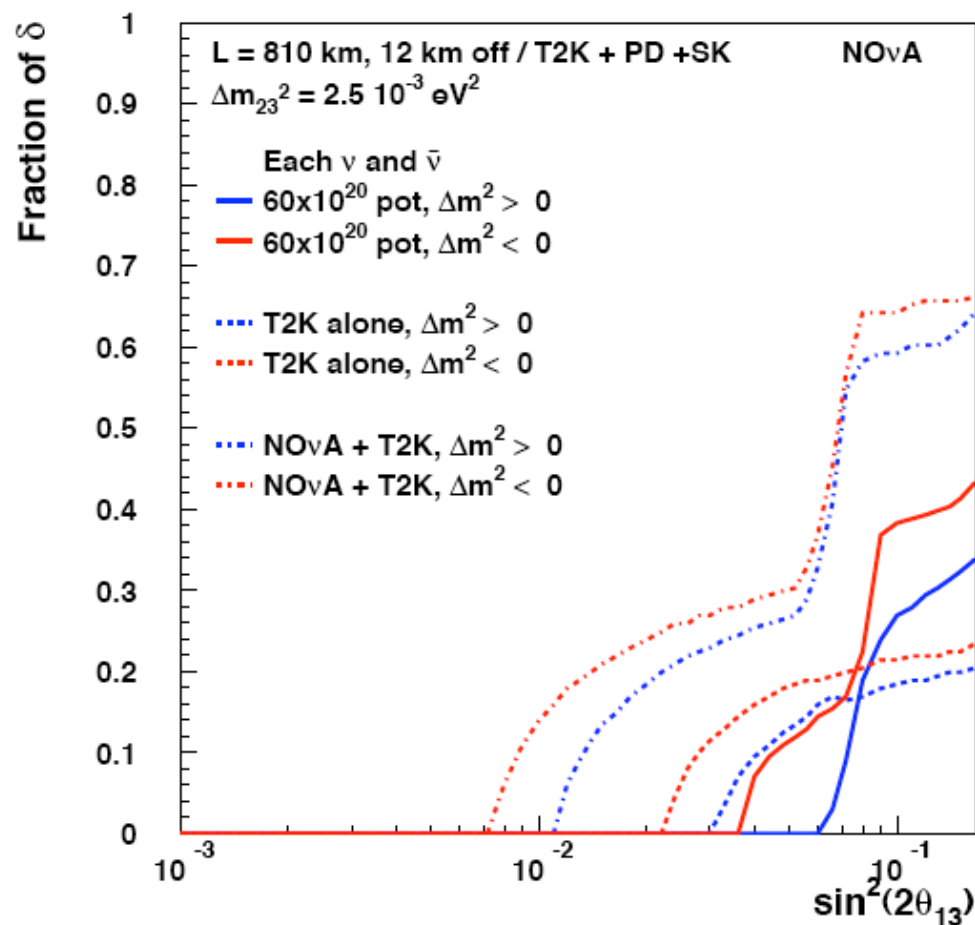
Mass Hierarchy Resolution Summary





3 σ Demonstration of CP Violation

3 σ Determination of CP Violation



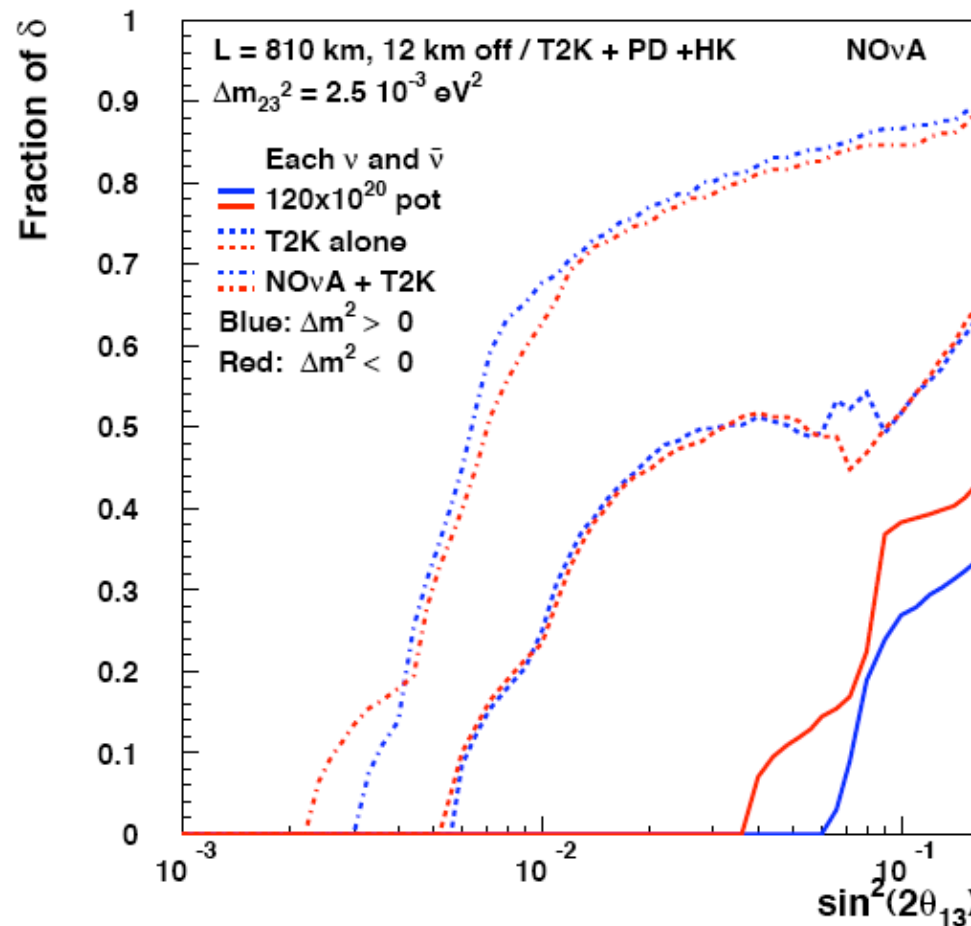
**With proton
drivers**

**(No 3 σ CP effect
in either T2K or
NOvA without
them.)**



3 σ Demonstration of CP Violation

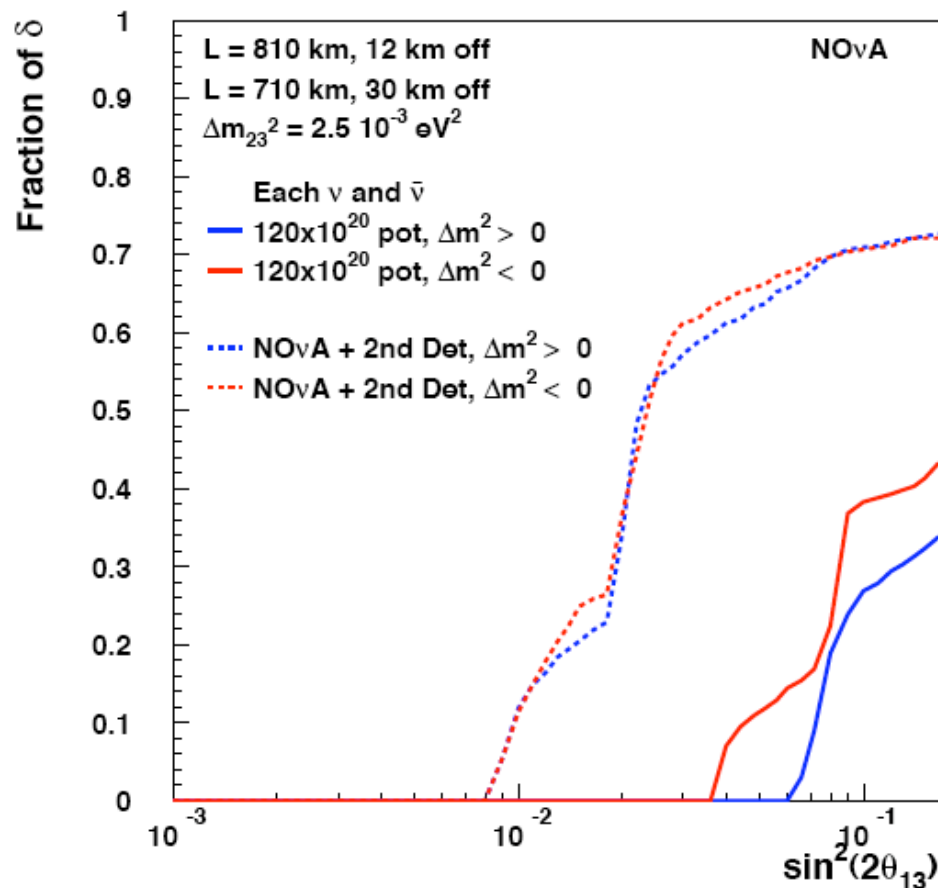
3 σ Determination of CP Violation





3 σ Demonstration of CP Violation

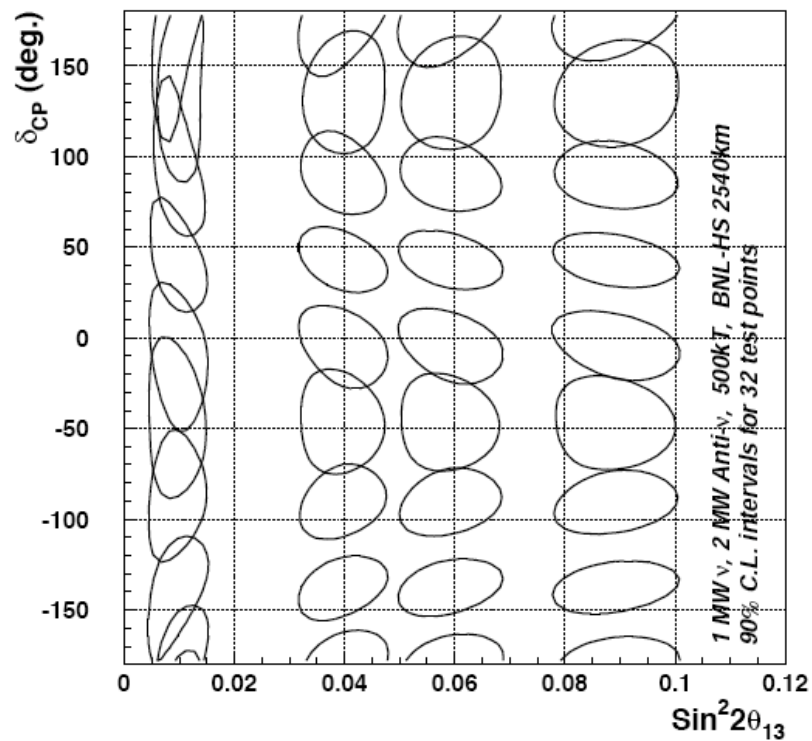
3 σ Determination of CP Violation



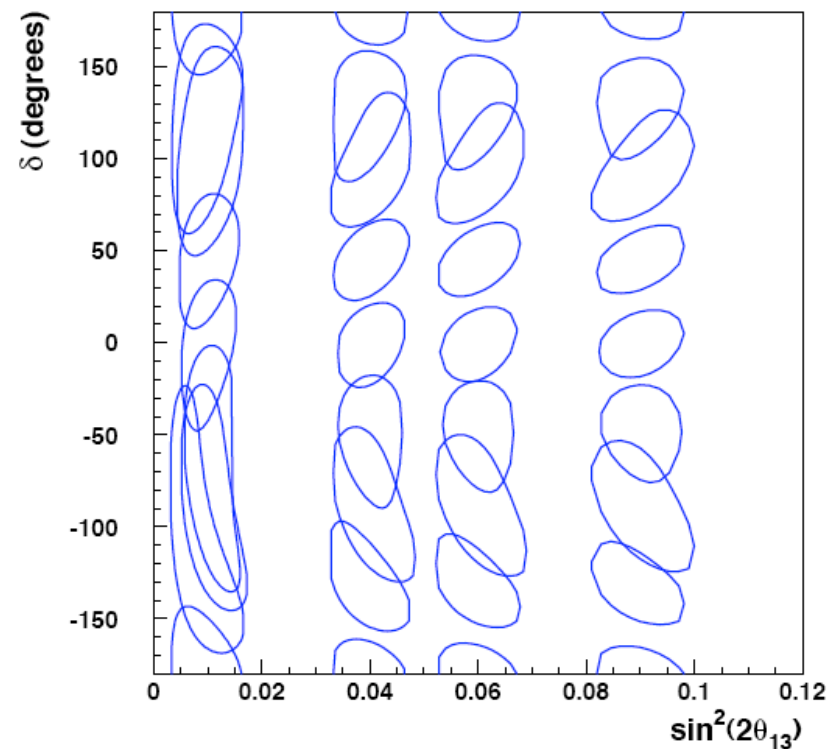
**2nd Off-axis
detector at the
2nd maximum**



90% CL δ vs. $\sin^2(2\theta_{13})$ 5 yrs each ν and anti- ν



BNL Proposal



NOvA + PD + 2nd Det



Conclusions to June PAC

- **NOvA provides a flexible approach to studying all of the parameters of neutrino oscillations**
 - A long baseline approach is crucial in the context of the world program.
 - NOvA is the first stage of a flexible program where each stage can be planned according to what has been learned in previous stages.
 - The range of the NOvA program is comparable to that of other conventional approaches.
 - NOvA can be started now (same scale as NuMI/MINOS).
 - The approval road is long. We need PAC approval now to keep NOvA and the Fermilab neutrino program timely.



PAC Response (1)

- “To establish a compelling physics case, NOvA must meet the following criteria:
 1. **Uniqueness.** NOvA must have a unique physics capability not achieved by any other experiments worldwide.
 2. **Competitiveness with T2K.** NOvA must compete with T2K, the Japanese program discussed above, within a similar time frame.
 3. **Competitiveness and/or complementarity with future experiments at reactors.** NOvA must compete in sensitivity with reactor experiments, or provide information not obtainable by reactor experiments.
 4. **Capability for evolution with a future neutrino program.** NOvA must allow a natural progression to CP violation studies with a future proton driver with the currently proposed detector at the same location.”



PAC Response (2)

- The report then summarizes the physics case and concludes “The Committee finds the proposal meets the above four criteria if the detector can be built in a timely manner.”
- “How soon must NOvA start taking data in order to be timely?... The Committee concludes that NOvA must start data-taking in the same time frame as T2K, and complete the far detector within four years to meet this criterion.... The Committee notes that the timely construction of NOvA is inconsistent with the present budget projection of the Laboratory.”



PAC Response (3)

- **“The Committee strongly endorses the physics case for the NOvA detector, and would like to see NOvA proceed on a fast track that maximizes its physics impact.”**
- **“The Committee encourages the Laboratory to work together with the funding agencies to put the necessary funding profile in place for a construction start in FY 2007, or in FY 2008 at the latest.”**
- **“The Committee strongly endorses the proposed R&D plan and urges the Laboratory to provide adequate support for timely completion of this program.”**



PAC Action

- **“While the Committee applauds [NOvA’s] progress, it concludes that Stage I approval at this time is premature. The collaboration should first complete the following steps:”**
 - **Complete the critical R&D**
 - Adequate signal from a full-length prototype
 - Cosmic ray background test
 - Engineering studies for TASD
 - **Make a final technology choice**
 - **Update the proposal**



R&D Activities

- Submitted a FY05 request to Fermilab for 513 k\$
- Acquiring materials for the vertical-slice test
 - 3rd iteration on the PVC extrusion
 - Mechanical and electronics design underway
- Design of the cosmic ray test progressing
 - Plan to use MINOS spare scintillator modules and electronics
- Mechanical studies of TAsD started
 - Half-scale prototype at Argonne using commercial extrusions (picture next slide)



Baseline Detector

50 kT

Liquid Scintillator:

1.2 m x 3 cm x 14.4 m

30-cell PCV extrusions,

24 extrusions/plane,

750 planes

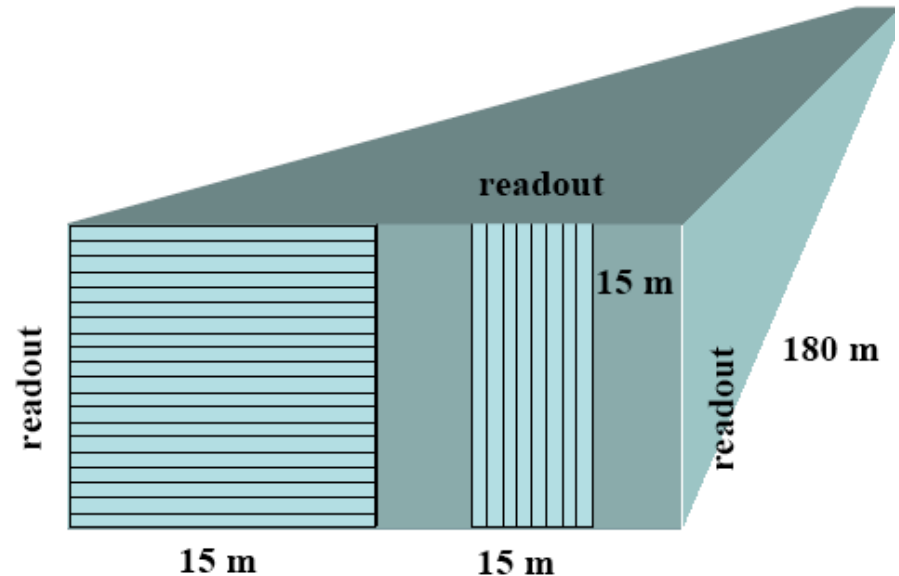
= 18,000 extrusions

= 540,000 channels

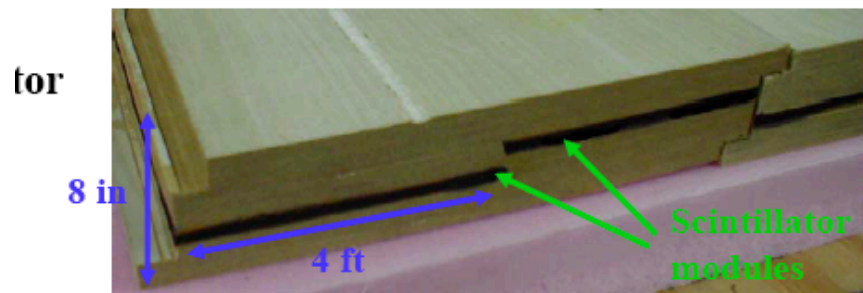
U-shaped WLS fiber into
APD readout

Absorber:

20 cm particleboard/
plane ($\sim 1/3 X_0$)



800 planes X, absorber, Y, absorber, ... = detector





Alternative Under Study: Totally Active Scintillator Detector (TASD)

25 kT

Liquid Scintillator:

1.28 m x 4.9 cm x 17.5 m

32-cell PCV extrusions,

14 extrusions/plane,

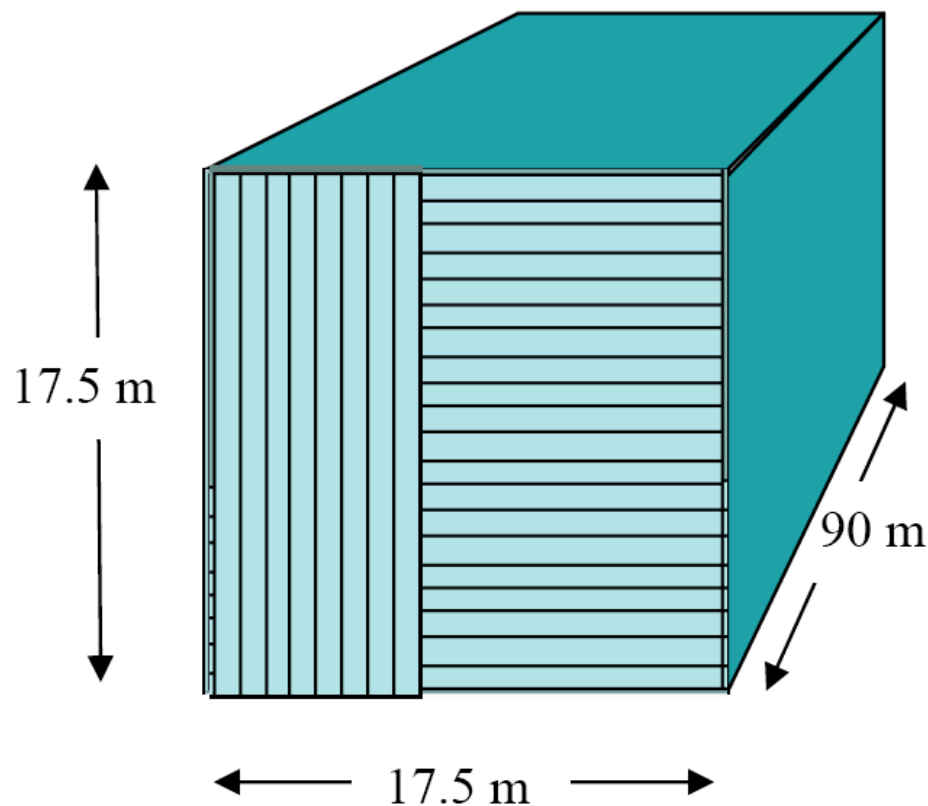
1845 planes

= 25,830 extrusions

= 826,560 channels

**U-shaped WLS fiber into
APD readout**

Absorber: None





Comparisons

	Baseline	TASD
Mass	50 kT	25 kT
Optimized ν_e efficiency	18%	32%
Optimized s/b	4.8	7.7
FoM	24.5	24
Cost	\$147M	\$159M

No obvious drawbacks to TASD so far.



Visit to the DoE

- On August 24, J. Cooper, D. Michael, H. Montgomery, S. Parke, S. Wojcicki, and I met with Robin Staffin, about 7 members of his staff, and Joe Dehmer for about 2 1/2 hours.
- I think we were successful in educating them on the physics case for NOvA and giving them a story they could sell in Washington. (Next talk)



Other Developments

- **APS Study endorsement. (Talk after next)**
- **Linear Collider technology decision.**
- **Proton Driver workshop (starts Wednesday).**
- **Designation of a new Fermilab director.**
- **National Academy study.**
- **A neutrino subpanel?**